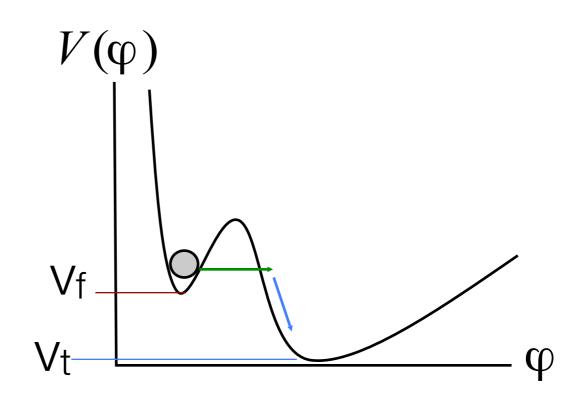
#### Can we see other universes?

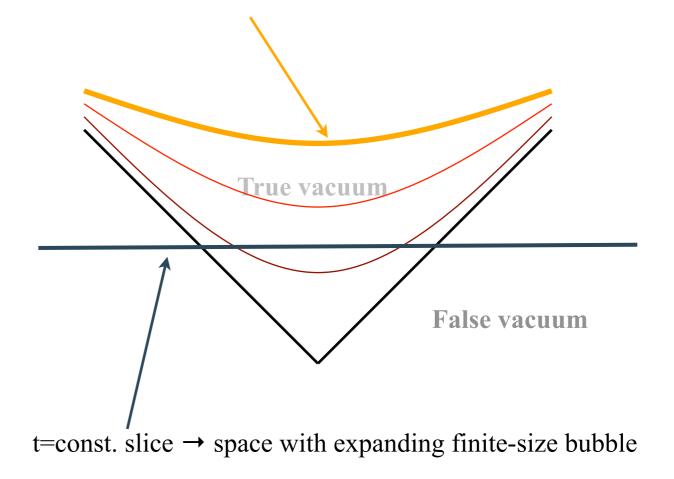
based on work by Anthony Aguirre, Matt Johnson & Assaf Shomer

(Multiple minima) +
 (slow transitions)
 = eternal inflation

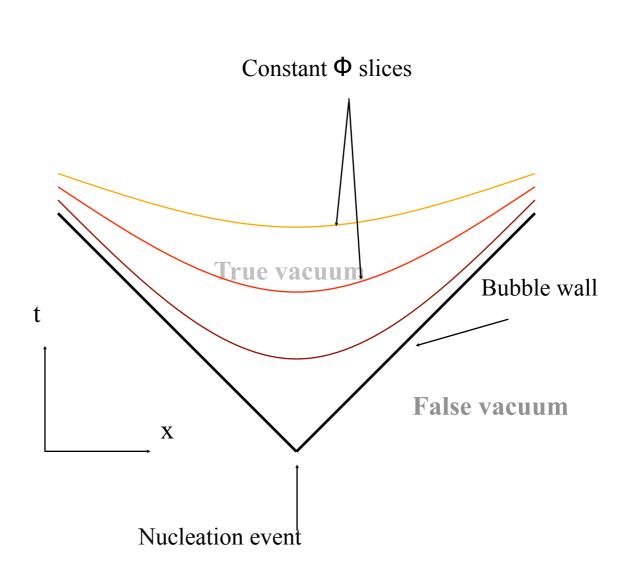


- (Multiple minima) +
  (slow transitions)
  = eternal inflation
- Each bubble has open FRW cosmology inside.

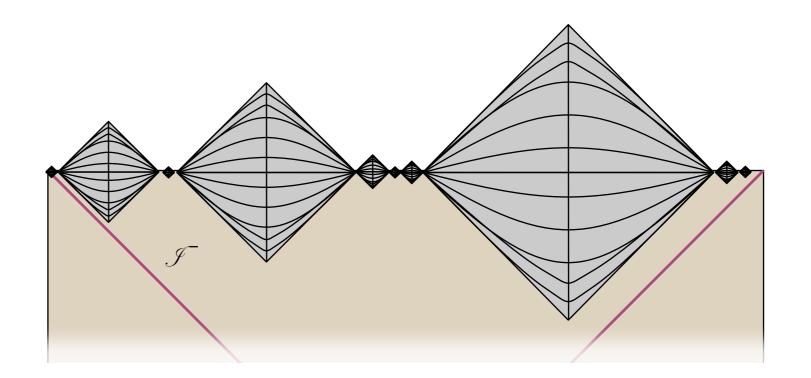




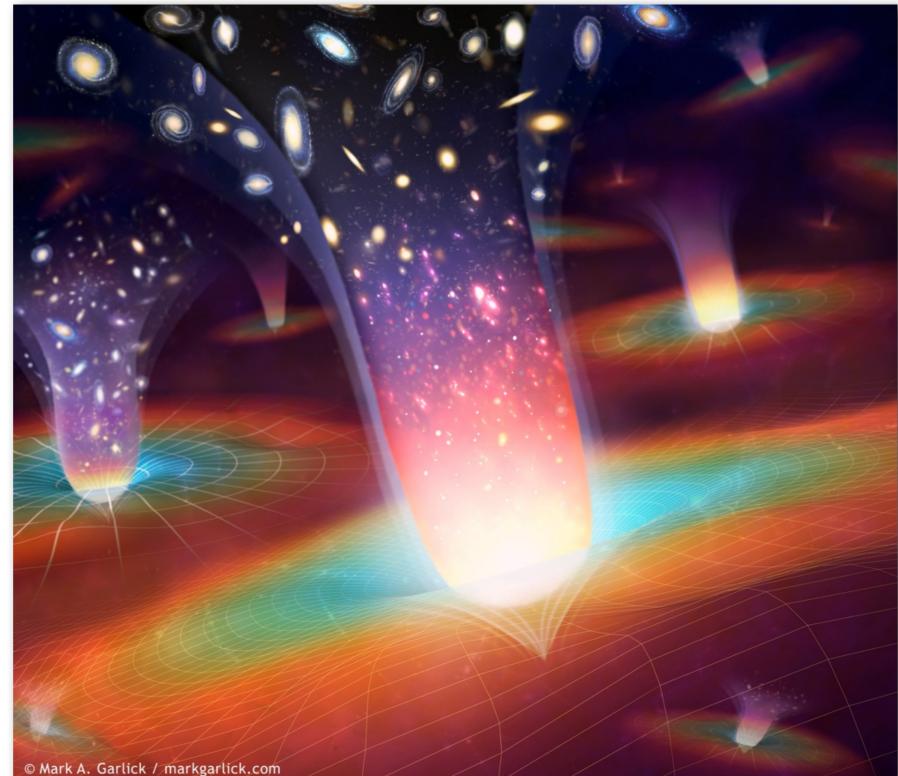
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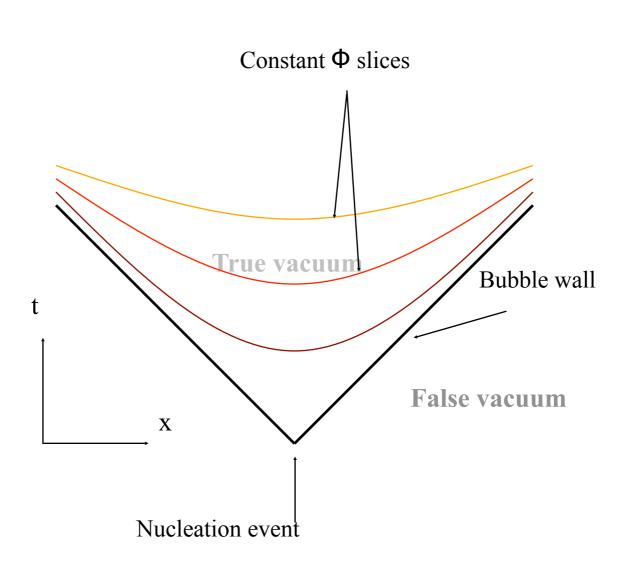
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- Infinitely many other bubbles, potentially different properties



- (Multiple minima) +
  (slow transitions)
  = eternal inflation
- Each bubble has open FRW cosmology inside.
- Infinitely many other bubbles, potentially different properties
- How do we test this picture?

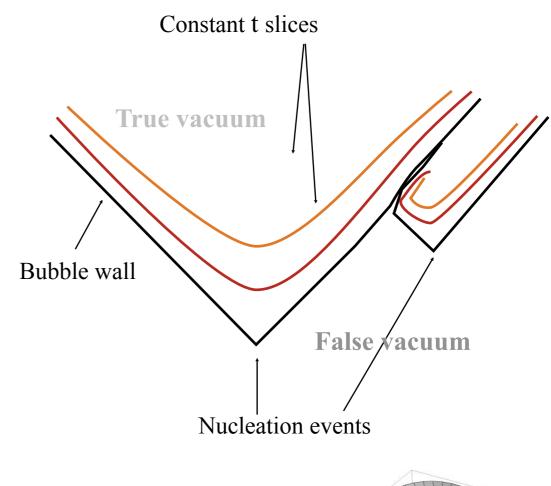


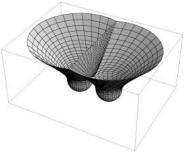
- Each bubble has open FRW cosmology inside.
- Can't get out.



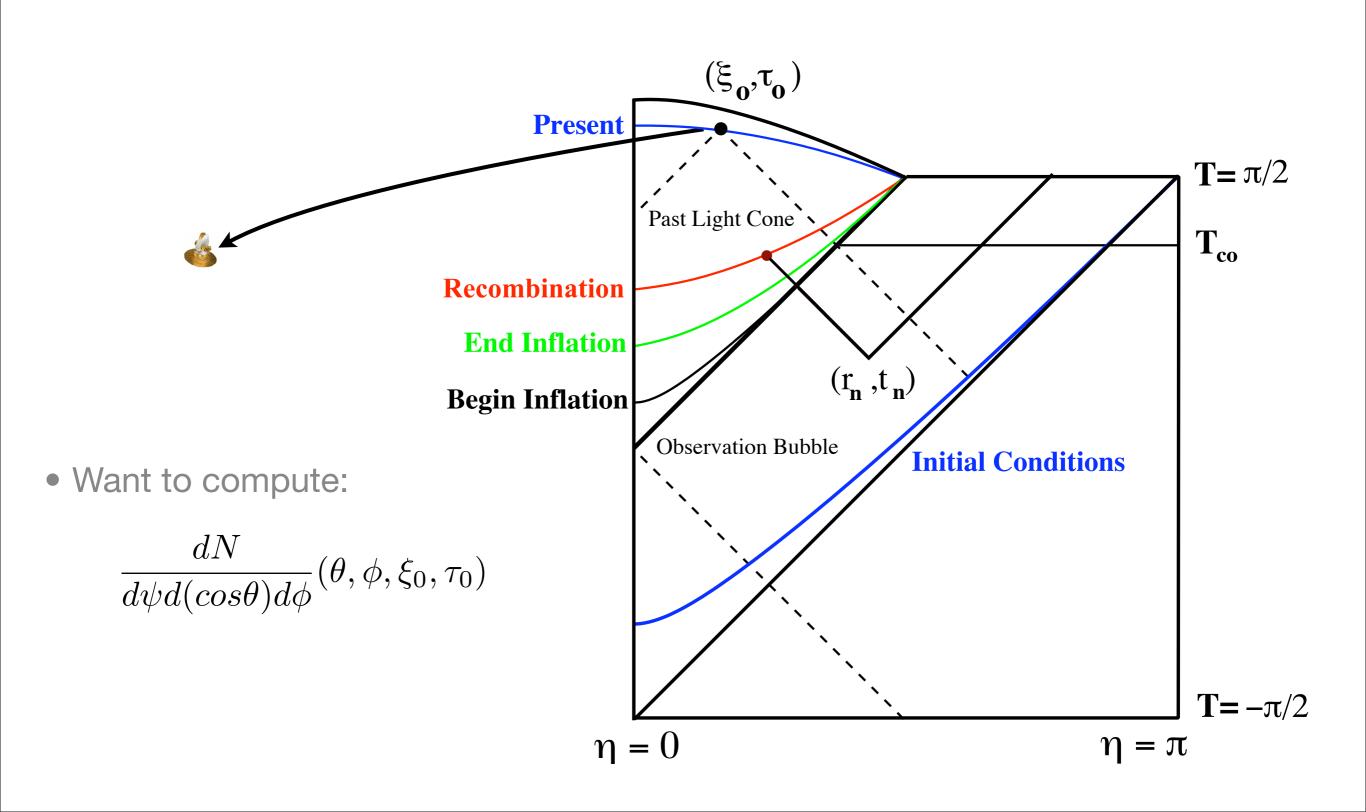
## Bubbles collide. Can we see the other ones?

- Probability: Most observers must have collisions in past.
- Suvivability: The collisions must not preclude the existence of the observers in question.
- Observability: The collision must be significant enough to be observable.

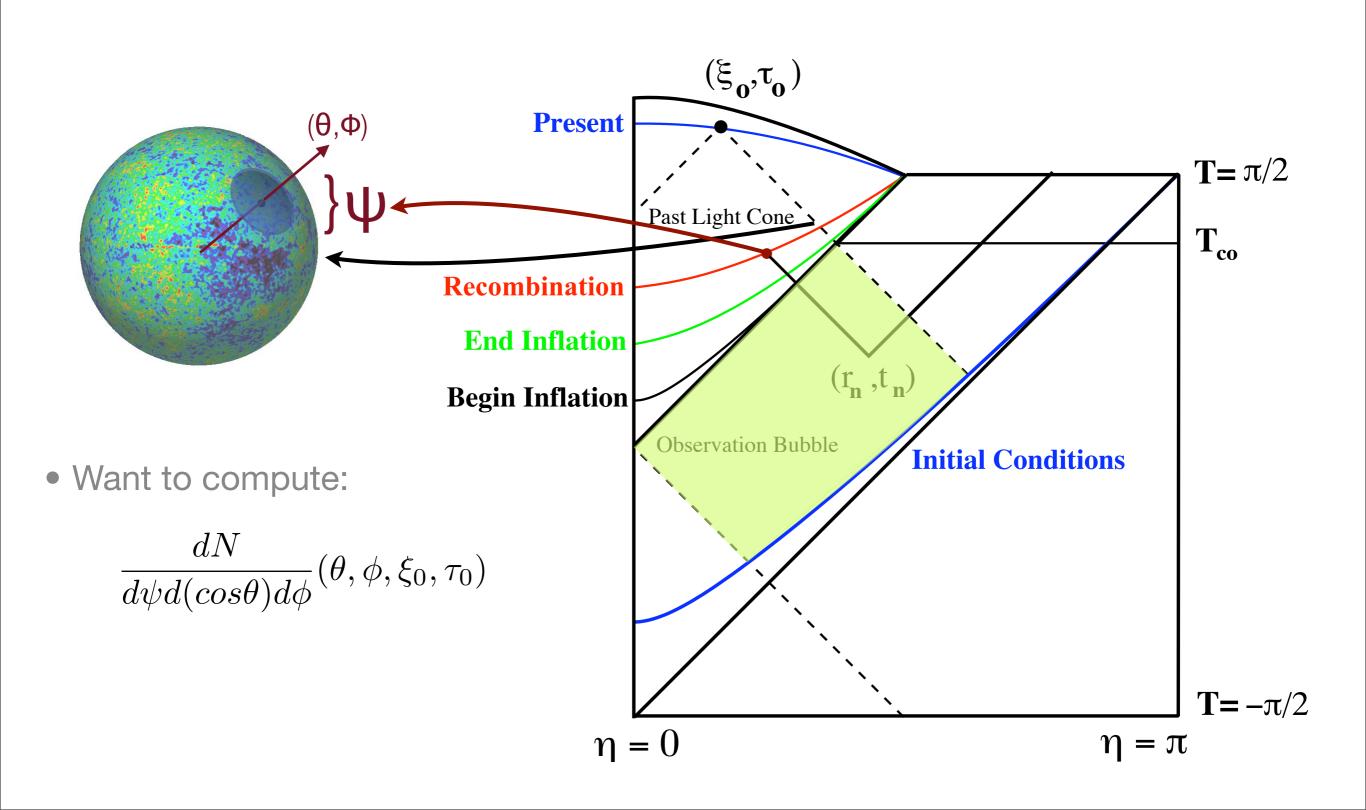




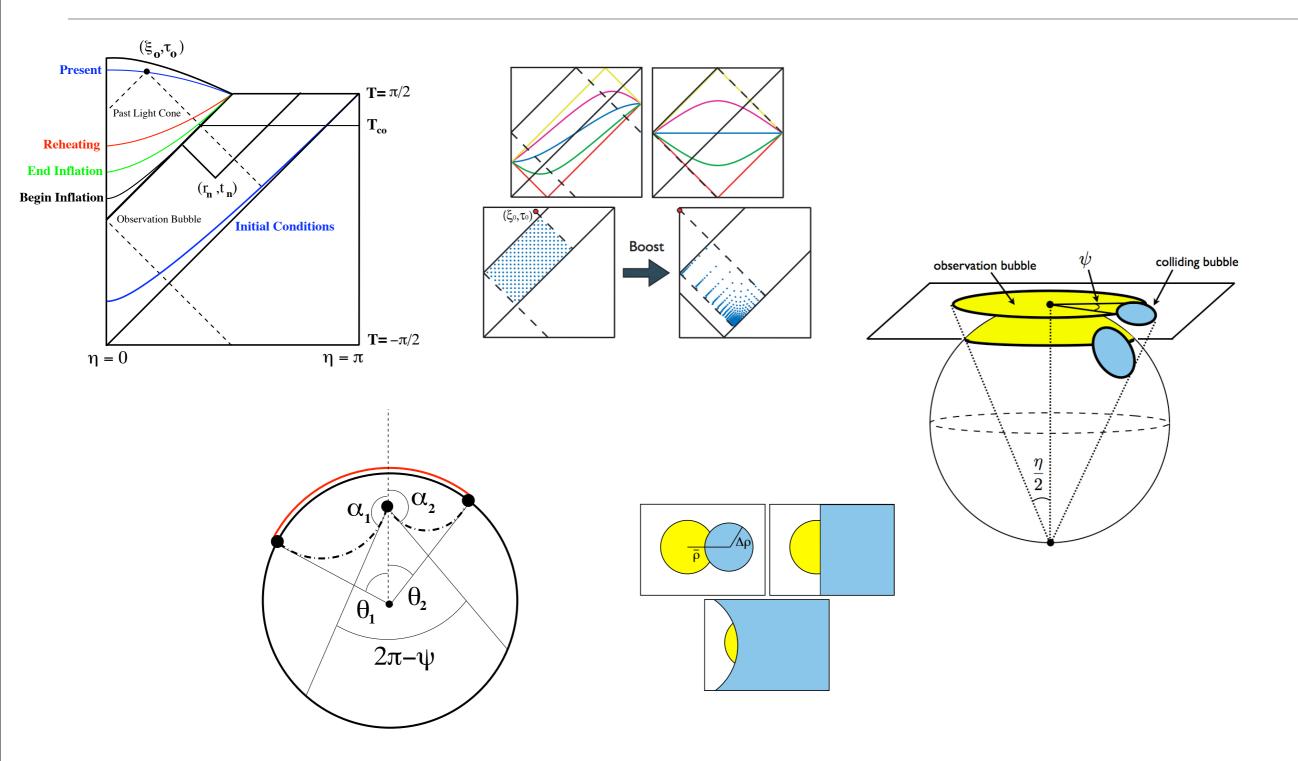
### The setup



### The setup

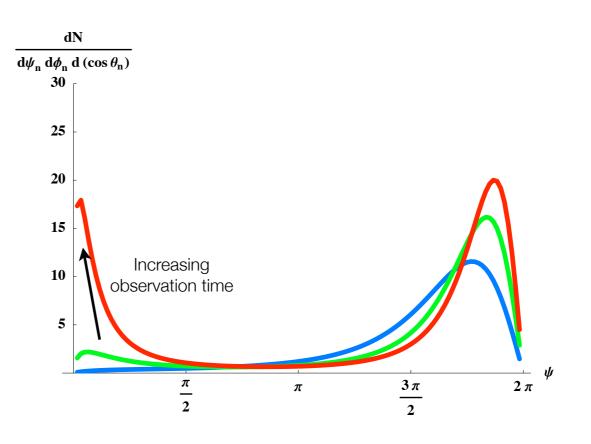


### Horrible geometry problem...



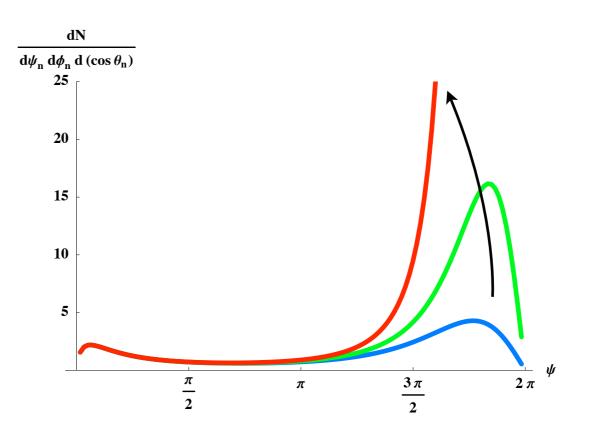
#### Results

- Two large-N regimes:
  - Late, small bubbles (as  $\tau_0 \rightarrow \infty$ , N  $\rightarrow \infty$  and  $\psi \rightarrow 0$ )
  - Early, large bubbles (as  $\xi_0 \rightarrow \infty$ , N  $\rightarrow \infty$  and  $\psi \rightarrow \infty$ )
- Anisotropic distribution (divergences where  $\theta = \theta_0$ )



#### Results

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## Implications

#### • Small late bubbles:

- Almost certainly 'perturbative'. ③
- Only seen if  $\lambda H_F^{-4} > (H_T/H_F)^2 \sim 10^{-100}$ .
- Essentially point sources, isotropically distributed.

#### • Large early bubbles:

- Seen at all but set of measure zero of points in open universe. 😳
- But not perturbative.

# Key open questions

- What does a general bubble collision (including inflation inside bubbles) look like?
- Can the observers with big observable collisions survive to observe them?
- Are there reasonable scenarios with small observable bubbles?
- How would bubbles appear in the CMB? What other consequences might they have? and/or, can any scenarios can be ruled out given the CMB we see?



